

## 1 Melting function

The noble gases concentration in the residual, melt-depleted mantle depends on the mineral-melt partition coefficient,  $D$ , and enters via the melt function,  $f = C_r/C_0$ , where  $C_r$  is the concentration in the residual mantle and  $C_0$  the concentration prior to melting.  $f$  is the fraction of noble gas left in the residual mantle, whereas the remainder is in the melt phase, which solidifies to form ocean crust. In our analysis it is assumed that all noble gases are lost from the ocean crust by the time it is subducted back into the mantle. We thus assume that a fraction  $f$  of the noble gases are lost to the Earth's oceans and atmosphere upon partial melting of a given mass of mantle. The relation between  $D$  and  $f$  is shown in Figure S1. Fractional and batch melting are end-member cases of the more general dynamic melting model with melt retention [32], where  $C_m$ , the concentration in the melt, is given by

$$C_m = C_{m,0} \left( \frac{D + \alpha - F(D + \alpha)}{D + \alpha - f_p(D + \alpha)} \right)^{((1-D)/(D+\alpha))}. \quad (1)$$

Here  $\alpha = f_p/(1 - f_p)$ ,  $f_p$  is retained melt fraction,  $C_0$  is the concentration in the solid prior to melting and  $C_{m,0}$  is defined as

$$C_{m,0} = C_0 (D + f_p(1 - D)). \quad (2)$$

The two end-member cases are fractional ( $f_p = 0$ ) and batch ( $f_p = F$ ) melting, where  $F$  is the melt fraction. For fractional melting

$$f = (1 - F)^{1/D-1} \quad (3)$$

and for batch melting

$$f = D(D + F - DF)^{-1}. \quad (4)$$

A reasonable upper limit for the fraction of retained melt is 2%, or  $f_p = 0.02$ . For our model calculations we therefore assume a melt retention of  $f_p = 0.02$  with partition coefficients of  $D_{\text{He}} = 0.0002$  and  $D_{\text{Ar}} = 0.0011$  [12]. Figure S1 shows that for  $D_{\text{He}} \approx 0.0002$  essentially all He is lost from the mantle residue upon melting and the same holds true for Argon (Ar) at  $D_{\text{Ar}} \approx 0.001$ . Although recent estimates for He mineral-melt partition coefficients differ by about one order of magnitude, high estimates of  $D_{\text{He}} = 0.0025 - 0.006$  [33] still result in at least 99% of He loss from the mantle under typical melting regimes, as shown in Figure S1. Because for any reasonable choice of  $D_{\text{He}}$ ,  $D_{\text{Ar}}$  and melt fraction  $F$ , the value  $f$  is small, our model results are not sensitive to the particular choice of  $f_p$  or  $D$ . This assertion is substantiated,

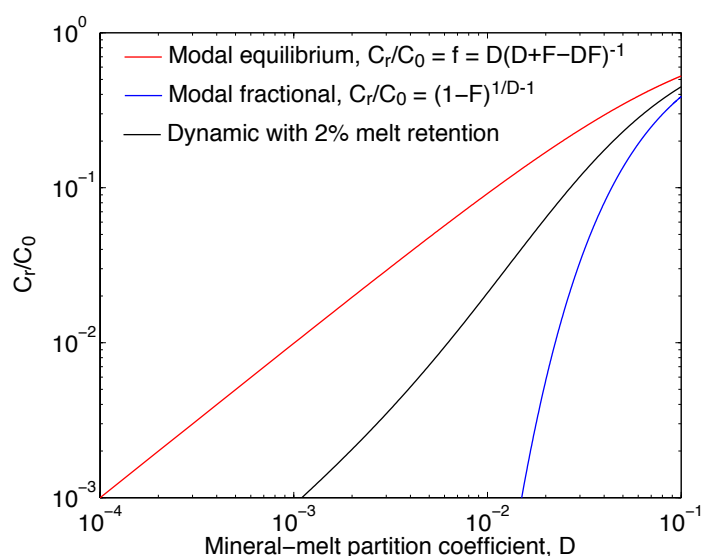


Figure S1:  $C_r/C_0$  for the residual mantle after partial melting, assuming  $F = 0.1$ .  $C_r/C_0$  is essentially zero for any reasonable values of  $D$ .

for example, for the case of instantaneous mixing given by Equation (1) and shown in Figure S2.

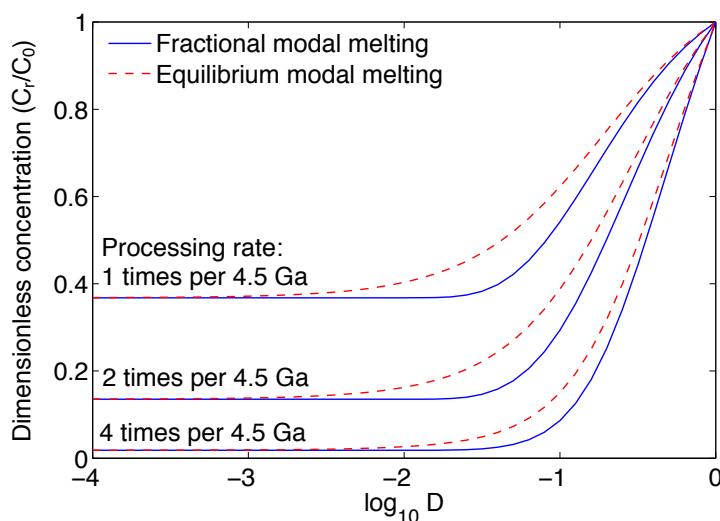


Figure S2:  $C_r/C_0$  after 4.5 Ga as a function of  $D$  for different values of  $N$ , the reservoir masses processed over 4.5 Ga, with and  $F = 0.1$  and assuming instantaneous mixing. Here,  $C_r$  is the concentration of a non-radiogenic, non-recycled noble gas in the mixed mantle reservoir after 4.5 Ga of mantle processing and  $C_0$  is the initial concentration. Note that the extent of outgassing asymptotically approaches  $D = 0$  and is essentially constant for  $D < 0.01$ .

## 2 Mass of unmixed recycled slab

Equation (7) in the Methods section gives the fraction of the mantle reservoir comprised by segregated (unmixed) recycled slab ( $M_S/M_0$ ). Figure (S3) shows the value of  $M_S/M_0$  as a function of reservoir masses processed over 4.5 Ga for different values of  $\tau$ . Unless a mantle reservoir is predominantly comprised of unmixed slab,  $\tau$  cannot exceed a few Ga (at least in a time-averaged sense), if the mantle reservoir has been processed at least once over Earth's history (i.e.,  $N \geq 1$ ).

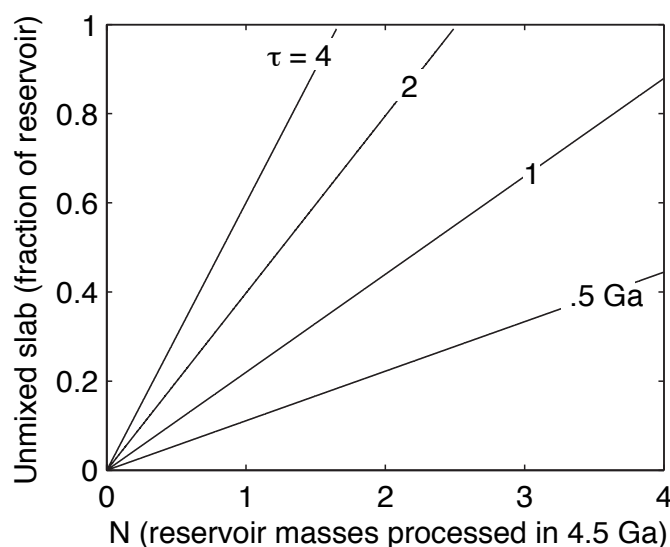


Figure S3: The fraction of segregated slab,  $M_S/M_0$  after 4.5 Ga as a function of  $N$  for different values of  $\tau$ .

## 3 Description of the geochemical reservoir model

We consider two geochemical mantle reservoirs denoted by subscripts  $U$  and  $L$ , with masses that are equivalent to the Earth's upper mantle and lower mantle, respectively. At time zero, the reservoirs consist entirely of a well-mixed region with BSE compositions. Recycling and mixing of the slabs into the mantle causes the upper and lower mantle reservoirs to evolve into a



1) to infer how He, Ar, U, Th, and K will partition between the oceanic crust and residual mantle. The plume melt flux is given by  $F Q_P$  and is enriched in He, Ar, U, Th and K. We assume that all noble gases are lost from the melt (ocean crust) and are not recycled. A fraction,  $x_C$  of U, Th, and K contained in the melt is assumed to be sequestered in the continental crust and thus not recycled to the mantle. The remainder is contained within the ocean crust. Neglecting the actual mass of continental crust produced over time, it is assumed that slab production, given by  $F Q_P + (1 - x_P)(1 - F) Q_P + Q_U$ , is balanced by slab subduction at a rate of  $Q_S$ . In other words, we assume that slabs are subducted at the same rate as they are produced, with the actual time lag between production and subduction implicitly accounted for by the characteristic mixing time,  $\tau$ . A fraction  $x_S$  of subducted slabs enters the lower mantle and mass balance requires that  $Q_L = x_S Q_S$ . Subducted slabs containing low concentrations of residual He, Ar, U, Th, and K mix with ambient mantle on a characteristic time scale,  $\tau_L$  and  $\tau_U$  for lower and upper mantle, respectively. We track the mass balance for  ${}^3\text{He}$ ,  ${}^4\text{He}$ ,  ${}^{40}\text{Ar}$ , U, Th and K in the mixed mantle reservoir and in the individual parts of unmixed slab by numerical integration.

To fully understand the range of permissible and realistic parameters and outcomes, we allow a number of parameters to vary independently over a large number of Monte Carlo simulations. The ranges of variable input parameters (summarized in Table S1) are:

- Present day processing rate  $Q_{L,present}$  falls within the range of present day plume flux estimates ( $2 \times 10^{13}$  to  $20 \times 10^{13}$  kg  $\text{y}^{-1}$  [27]).
- Present day processing rate  $Q_{U,present}$  is equal to present day upper-mantle processing rate through mid-ocean ridges ( $7 \times 14$  kg  $\text{y}^{-1}$  [34]).
- $Q_L$  is “anchored” at present day values and may vary through time as  $Q_L(t) = Q_{L,present} \exp(t/4.5 \text{ Ga})$ , so that the entire mass of lower mantle processed over 4.5 Ga ranges from 0.0304 (present day low estimate of plume flux) to 2 lower mantle masses.
- $Q_U$  is “anchored” at present day values and may vary through time as  $Q_U(t) = Q_{U,present} \exp(\alpha t/4.5 \text{ Ga})$ , so that the entire mass of upper mantle processed over 4.5 Ga ranges from 2.6 (present day upper-mantle processing rate through mid-ocean ridges) to 10 upper mantle masses.

- The residual plume mantle that mixes with upper mantle,  $x_P$ , is allowed to vary from 0 to 1, but assumed constant for any given model case.
- The fraction of slabs that subduct into lower mantle is  $x_S$ . The value of  $x_S$  is allowed to vary randomly for each case, but is always constrained to a value such that  $x_S Q_S \geq Q_P$ .
- The fraction of U, Th and K that is extracted and stored in the continental crust,  $x_C$ , is allowed to vary between 0 and 1 and is given by  $x_C(t) = r_1 * \exp(-10^{r_2} t / 4.5 \text{ Ga})$ , where  $r_1$  is a random number between 0 and 1 and  $r_2$  is a random number between -5 and 5. Both  $r_1$  and  $r_2$  are independent and differ for each Monte Carlo realization.
- The initial (primordial) concentration of  $^3\text{He}$  in upper and lower mantle falls between  $0.75 \times 10^{11}$  to  $1.5 \times 10^{11}$  atoms  $\text{g}^{-1}$ .
- Characteristic upper and lower mantle mixing times may vary independently within the range 0-4 Ga.

We thus examine a broad range of parameter space to fully explore the realm of feasible mantle evolution models for a coupled upper-lower mantle system. Constant initial conditions for all simulations are as follows:

- Initial (primordial)  $^3\text{He}/^4\text{He} = 120R_A$  for the entire mantle.
- Initial (primordial)  $^{40}\text{Ar} = 0$  for the entire mantle.
- Initial (primordial) mantle concentrations for radioactive elements are based on Bulk Silicate Earth estimates (VanSchmus, 1995) and are  $^{238}\text{U} = 40 \text{ ppb}$ ,  $^{235}\text{U} = 12 \text{ ppb}$ ,  $^{232}\text{Th} = 95 \text{ ppb}$  and  $^{40}\text{K} = 327 \text{ ppb}$ .

From the Monte Carlo realizations we select outcomes that each meet all of the following criteria:

- Present day  $^3\text{He}/^4\text{He}$  in the upper mantle =  $8.5 \pm 0.5R_A$ .
- Present day U content of the upper mantle falls within the range of estimated values by Workman and Hart [28] (3.2 ppb) and by Boyet and Carlson [29] (5.4 ppb).
- Present day K content of the upper mantle falls within the range of estimated values by Workman and Hart [28] (50 ppm) and by Boyet and Carlson [29] (68.4 ppm).
- The present day rate of  $^3\text{He}$  loss from the upper mantle falls within 75-100% of the estimated  $^3\text{He}$  flux to the oceans (1060 mole  $\text{y}^{-1}$  [30]).
- Present day  $^{40}\text{Ar}$  in the mantle is within 40-60% of radiogenic  $^{40}\text{Ar}$  produced over 4.5 Ga within a bulk-silicate Earth (BSE) mantle, assuming  $K_{BSE} = 235 \text{ ppm}$  and zero primordial  $^{40}\text{Ar}$ .

Table S1: Parameters used in the geochemical mantle reservoir model

Parameter	Description	Value
$M_U$	mass of upper mantle	$1.1 \times 10^{24}$ kg
$M_L$	mass of lower mantle	$2.96 \times 10^{24}$ kg
$Q$	instantaneous processing rate	$\lesssim 10^{16}$ kg $y^{-1}$
$Q_{U,present}$	present day processing rate	$7 \times 10^{14}$ kg $y^{-1}$
$Q_{L,present}$	present day processing rate	$2 - 20 \times 10^{13}$ kg $y^{-1}$
$\tau$	characteristic mixing time	0-4 Gy
$(^3\text{He}/^4\text{He})_{t=0}$	initial ratio	120 $R_A$
$^3\text{He}_{t=0}$	initial $^3\text{He}$ concentration	$0.75 - 1.5 \times 10^{11}$ atoms $g^{-1}$
$^{40}\text{Ar}_{t=0}$	initial $^{40}\text{Ar}$ concentration	0 ppb
$^{238}\text{U}_{t=0}$	initial $^{238}\text{U}$ concentration	40 ppb [35]
$^{235}\text{U}_{t=0}$	initial $^{235}\text{U}$ concentration	12 ppb [35]
$^{232}\text{Th}_{t=0}$	initial $^{232}\text{Th}$ concentration	95 ppb [35]
$^{40}\text{K}_{t=0}$	initial $^{40}\text{K}$ concentration	327 ppb [35]
$D_{\text{He}}$	mineral-melt partition coefficient	$2.0 \times 10^{-4}$ [12]
$D_{\text{Ar}}$	mineral-melt partition coefficient	$1.1 \times 10^{-3}$ [12]
$D_U$	mineral-melt partition coefficient	$3.0 \times 10^{-3}$ [36]
$D_{\text{Th}}$	mineral-melt partition coefficient	$3.4 \times 10^{-3}$ [36]
$D_K$	mineral-melt partition coefficient	$2.2 \times 10^{-3}$ [36]
$x_P$	frac. residual plume into upper mantle	0 to 1
$Q_S$	mass flux of subducting slab	$Q_S = Q_L + Q_U$
$Q_P$	plume flux	$\leq Q_L$
$x_S$	fraction slab into lower mantle	$x_S Q_S \geq Q_P$
$x_C$	fraction U, Th, K to continental crust	$x_C(t) = r_1 * \exp(-10^{r_2} \hat{t})$
$\hat{t}$	dimensionless time	$t/(4.5 \times 10^9)$
$r_1$	random number	0 to 1
$r_2$	random number	-5 to 5

## 4 Depletion of the mantle in Samarium and Neodymium

To demonstrate that our model results for noble gases are also consistent with the depleted character of the mixed region of the upper and lower mantle, we present model predictions for the Neodymium (Nd) isotopic evolution (Figure S5). The approach taken for the Samarium-Neodymium isotopic system is identical to our modeling of the radioactive elements U, Th and K. We use initial (bulk silicate Earth) abundances of 1.25 ppm for Nd and 0.406 ppm for Sm [37], as well as partition coefficients of 0.031 and 0.045 for Nd and Sm, respectively [28]. Although we do not model recycling of continental crust explicitly, our model results imply that recycled continental crust in the mantle may comprise approximately 0.5% of the mantle. For model cases that meet all aforementioned noble-gas related constraints, as well as  $9 < \varepsilon^{143}\text{Nd} < 10$  for the mixed region of the upper mantle, we find that the mixed lower mantle assemblage with the primitive noble gas signatures ( $^3\text{He}/^4\text{He}$  ratio between 35-40  $R_A$ ) has present day  $\varepsilon^{143}\text{Nd}$  between 2 and 5. This is within the range for FOZO-A and FOZO-B [38].



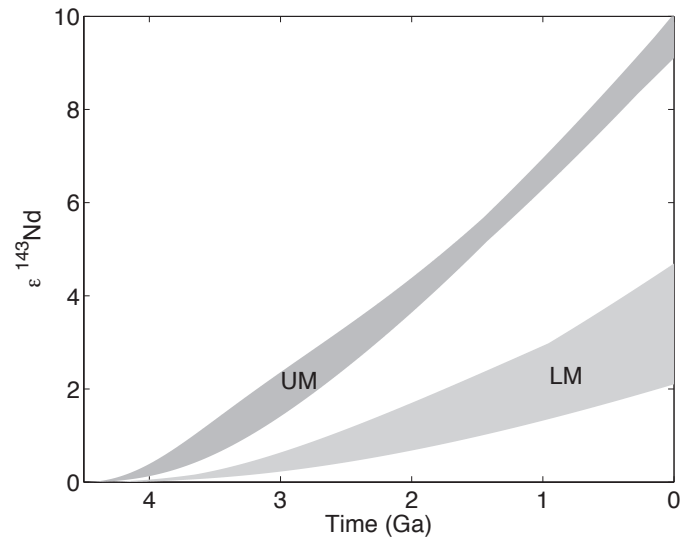


Figure S5:  $\epsilon^{143}\text{Nd}$ , defined as  $([^{143}\text{Nd}/^{144}\text{Nd}]_{\text{sample}} / [^{143}\text{Nd}/^{144}\text{Nd}]_{\text{CHUR}} - 1) \times 10^4$ , where CHUR is the Chondritic Uniform Reservoir [39], as a function of time. Shaded areas represent the predicted range of  $\epsilon^{143}\text{Nd}$  for the mixed portion of the lower mantle (LM) for those model cases that meet all of the aforementioned noble gas observational constraints (i.e., cases shown in Figure 4 of the text) and for which with  $9 < \epsilon^{143}\text{Nd} < 10$  in the mixed portion of the upper mantle (UM).

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